

# A Song of D-branes and Fluxes

Wieland Staessens (JdC)

*based on* [1807.00620](#), [1807.00888](#) ([1503.01015](#), [1503.02965](#) [*hep-th*])

*with G. Shiu*



Instituto de Física Teórica  
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Madrid



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Universidad Autónoma  
de Madrid

# Inflation in Type IIA

Guth, Linde, Mukhanov, Steinhardt, Starobinsky,...

- Inflationary epoch = cure for horizon problem and flatness problem
- nearly scale invariant, nearly Gaussian CMB data:

$$n_s - 1 = 2\eta - 6\epsilon, \quad r = 16\epsilon$$

in agreement with slow-roll single scalar field  $w$ / potential  $V$

$$\epsilon \equiv \frac{M_{Pl}^2}{2} \left( \frac{V'}{V} \right)^2 \ll 1, \quad |\eta| \equiv \left| M_{Pl}^2 \frac{V''}{V} \right| \ll 1 \quad \text{during inflation}$$

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reviews: Baumann (2009); Baumann-McAllister (2009,2014); Westphal (2014); ...

- Inflation in String Theory tied with Moduli Stabilization

D3/7-brane position moduli [Burgess et al \('01\)](#), [Dvali et al \('01\)](#), [Dasgupta et al \('02\)](#)  
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## Type IIB

Kähler Moduli (e.g. Fibre) [Cicoli-Burgess-Quevedo \('08\)](#) ([Cicoli](#), [Shukla](#))  
 Kähler Axions (aligned natural, N-flation, monodromy, kinetic alignment)

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## Type IIA

<u>Inflaton</u>	<u>Source</u>	<u>Potential</u>
Volume $\rho$	NS-flux & RR-flux	$\rho^{-3,3-p}$
Dilaton $s$	NS-flux & RR-flux & O6/D6	$s^{-2,-3,-4}$
$B_2$ -axion	RR-flux	$b^{1,2,3}$
$C_3$ -axion		

- Note: flux stabilization  $\rightsquigarrow$  only linear combination  $C_3$ -axions stabilized  
[DeWolfe-Giryavets-Kachru-Taylor \(2005\)](#), [Cámara-Font-Ibáñez \(2005\)](#)

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# Stringy Axions & Effective Decay Constant

w/ Shiu-Ye 1503.01015, 1503.02965 [hep-th]

- Type II String Theory compactifications Blumenhagen-Körs-Lüst-Stieberger ('06); Ibáñez-Uranga ('12)  
 $\rightsquigarrow$  Closed string axions  $a^i$  from dim. red. of  $p$ -forms  $C_{(p)}$  on  $\mathcal{M}_{1,3} \times \mathcal{X}_6/\Omega\mathcal{R}$   
 $(C_{(p)} \in \text{RR-forms} + \text{NS 2-form in Type II})$

$$a^i \equiv (2\pi)^{-1} \int_{\Sigma^i} C_{(p)}, \quad p\text{-cycle } \Sigma^i \subset \mathcal{X}_6, \quad i \in \{1, \dots, \left. \begin{matrix} h_{11} \\ h_{21} + 1 \end{matrix} \right\}$$

- Type II String Theory compactifications w/ D-branes  
 $\rightsquigarrow$  4d EFT with mixing axions + fermions (anomaly cancellation) (Dudas' talk)  
Aldazabel-Franco-Ibáñez-Rábadan-Uranga ('01)

$$S_{axion}^{\text{eff}} = \int \left[ \frac{1}{2} \sum_{i,j=1}^N \underset{\substack{\text{metric} \\ \text{mixing}}}{\mathcal{G}_{ij}} (da^i - k^i A) \wedge \star_4 (da^j - k^j A) - \frac{1}{8\pi^2} \left( \sum_{i=1}^N r_i a^i \right) \text{Tr}(G \wedge G) + \mathcal{L}_{gauge} + \mathcal{L}_\psi \right]$$

↙  $U(1)$  mixing  $k^i \neq 0$ 
↘ anomalous coupling

- Diagonalisation of kinetic and potential terms  
 $\Rightarrow$  effective decay constant  $f_{\text{eff}}$  with moduli dependence



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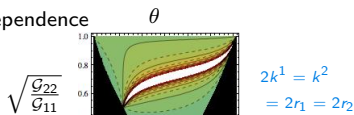
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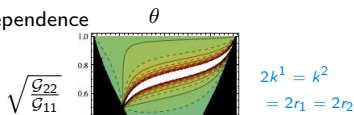
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# From UV (Fermions) to IR (Infladrons)

Shiu-W.S. (1807.00620, 1807.00888)

- Integrating out  $U(1) \rightsquigarrow 4\psi$  interactions (N-JL):  $\frac{g_L g_R}{2M_{\text{St}}^2} [(\bar{\psi}\psi)^2 + (\bar{\psi}i\gamma^5\psi)^2]$
- $\theta$ -vacuum of  $SU(N)$  YM breaks  $U(1)$  explicitly
  - ★ Instanton-induced effective fermion interactions 't Hooft ('76) Callan-Dashen-Gross ('78)

$$\mathcal{L}'_{\text{'t Hooft}} = C e^{-\frac{8\pi^2}{g^2} + i\theta} \det(\bar{\psi}_L \psi_R) + h.c.$$

at strong coupling for  $SU(N) \rightsquigarrow$  effective fermion mass

- ★ Fermion Confinement  $\Rightarrow$  Fermion condensate  $(\langle \bar{\psi}_L \psi_R \rangle_\theta \neq 0)$  Casher (1979)  
 $4\psi$  interactions  $\rightsquigarrow$  fermion mass  $M \sim -\frac{1}{M_{\text{St}}^2} \langle \bar{\psi}_L \psi_R \rangle_\theta$
- $E < \Lambda_5$ : bound state  $\bar{\psi}\psi \rightarrow$  EFT for composite scalar  $\Phi(x) = \sigma(x)e^{i\frac{\eta}{f}}$   
 Weinberg ('79)
- mass spectrum in vacuum

$$\begin{array}{c} f_\xi \ll f \\ \downarrow \\ m_\eta < m_\sigma \ll m_\xi \end{array}$$

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$$V = -\mu^2 \Phi^\dagger \Phi + \frac{\lambda}{2} (\Phi^\dagger \Phi)^2 + \Lambda_s^2 \left( \kappa e^{i\frac{\xi}{f_\xi} + i\theta} \det(\Phi) + \kappa e^{-i\frac{\xi}{f_\xi} - i\theta} \det(\Phi^\dagger) + M\Phi + M\Phi^\dagger \right)$$

Set by the  $U(1)$  symmetries in the model with spurions  $e^{i\theta}$  and  $M$

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- mass spectrum in vacuum massive  $(\sigma, \eta) =$  **INFLADRONS**

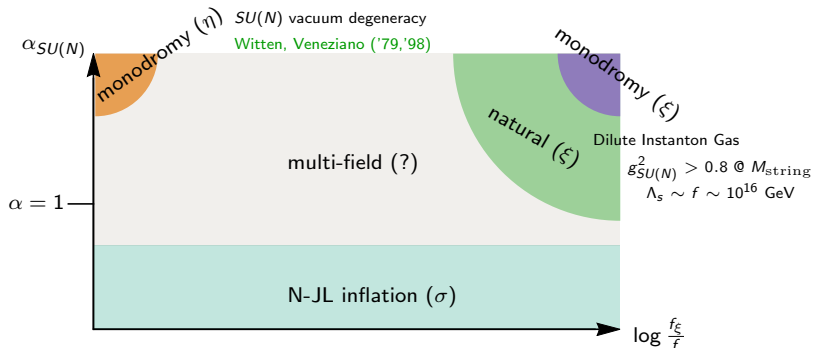
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# Phases of Axion Inflation

Shiu-W.S. (1807.00888)



$\sim$  Starobinsky inflation

strongly coupled  $4\psi$  interactions

see talk @ StringPheno 2016

see also Inagaki-Odintsov-Sakamoto ('15-'17)



# Natural-like Inflation

Shiu-W.S. (1807.00620, 1807.00888)

$\xi =$  inflaton candidate

with  $f \ll f_\xi$  and  $m_\xi \ll m_\eta < m_\sigma$

Viable inflationary model requires control over corrections:

- (1) perturbative QFT corrections constrained by perturbative  $U(1)$  symmetry  
Weinberg ('79), Coleman-Weinberg ('73), Hill-Salopek ('92)
- (2) back-reaction of heavy infladrons on inflationary trajectory  
see e.g. Stewart ('94), Lazarides-Panagiotakopoulos ('95), Lyth-Stewart ('96),  
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- (3) Pert. & Non-pert. gravitational corrections

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- (3) Pert. & Non-pert. gravitational corrections

# Natural-like Inflation

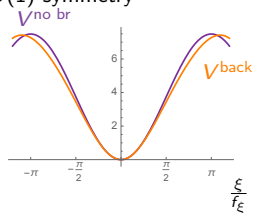
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with  $f \ll f_\xi$  and  $m_\xi \ll m_\eta < m_\sigma$

Viable inflationary model requires control over corrections:

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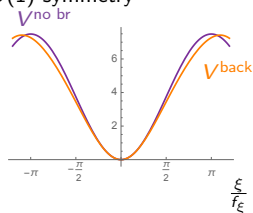
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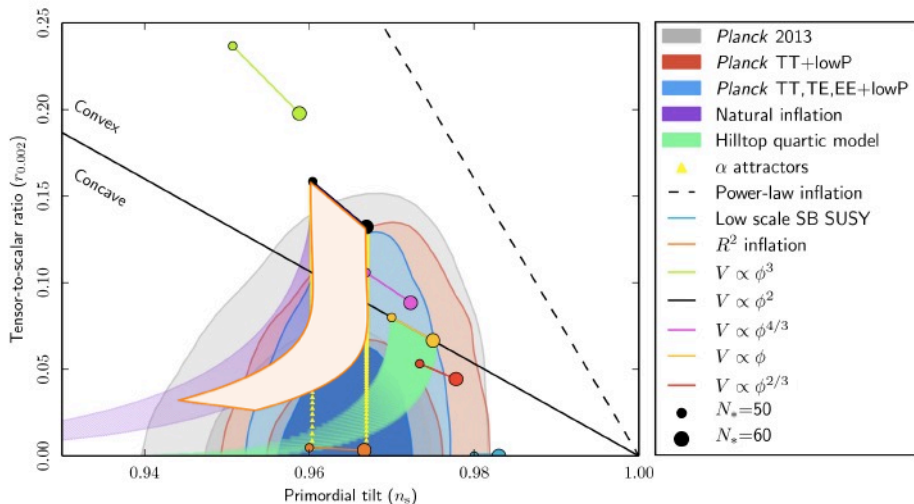
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based on PLANCK A&A 594, A20 (2016)

# Constraints from Gravity

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- Axionic wormholes with metric  $ds^2 = dr^2 + a^2(r)d\Omega_3^2$  and  $a(0) > M_{\text{st}}^{-1}$

$$S_{\text{AW}} = \int \sqrt{g_E} \left[ -\frac{M_{\text{Pl}}^2}{2} R_E - \underbrace{\frac{f_\xi^2}{2} g_E^{mn} \partial_m \xi \partial_n \xi}_{\text{Giddings-Strominger ('88)}} + \underbrace{\frac{1}{2} g_E^{mn} \partial_m \sigma \partial_n \sigma - \frac{\sigma^2}{2} g_E^{mn} \partial_m \eta \partial_n \eta + V_{\text{per}}(\sigma)}_{\text{Abbott-Wise ('89)}} \right]$$

axion charge  $w_\xi \gg 1$

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↓

$$S_{\text{GS}} \sim w_\xi \frac{M_{\text{Pl}}}{f_\xi} \gg 1$$

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Gravit. Instanton corrections highly suppressed Montero-Uranga-Valenzuela ('15)  
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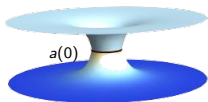
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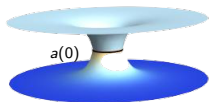
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# Conclusions and Outlook

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- Rich UV theory with mixing axions, gauge dof and fermions  
 $\rightsquigarrow$  rich IR theories in terms of axions and infladrons
- $\neq$  phase of gauge theories  $\rightsquigarrow \exists \neq$  inflationary models (natural, monodromy, Starobinsky)
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## Open issues

 Full String Theory construction including moduli stabilisation

Conlon (2006), Cicoli-Dutta-Maharana (2014), Blumenhagen-(Font-Fuchs-)Herschmann-Plauschinn(-Sekiguchi-Wolf) (2014/15),...

 Verification of WGC and other swampland conjectures

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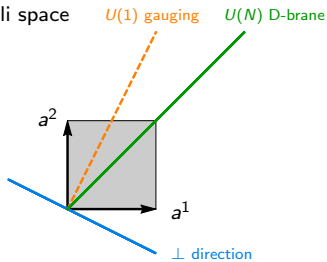
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Dziękuję ci bardzo

Thank you very much

# Effective Action & Effective Decay Constant

- Geometric picture of axion moduli space



- different from  $N$ -enhancement mechanisms:  $f_{\text{eff}} \sim N^p f$  with  $p \geq 1/2$ ,  
 Dimopoulos-Kachru-McGreevy-Wacker (2005), Choi-Kim-Yung (2014), Bachlechner-Long-McAllister (2014/15), Junghans (2015)

# Chiral Symmetry Breaking & Mass Generation

- Global  $U(1)$  symmetry  $J_{U(1)}^\mu = q_+ \bar{\psi} \gamma^\mu \psi + q_- \bar{\psi} \gamma^\mu \gamma^5 \psi$  is broken by gauge instantons & Yukawa-coupling ( $4\psi$ -coupling) in  $\theta$ -vacuum

$$\partial_\mu J_{U(1)}^\mu = -q_- \frac{1}{16\pi^2} \text{Tr}(\varepsilon^{\mu\nu\rho\sigma} G_{\mu\nu} G_{\rho\sigma}) + 2q_- M \bar{\psi} i \gamma^5 \psi$$

- Shift symmetry  $\xi \rightarrow \xi + \varepsilon_\xi$  is broken by gauge instantons

$$\partial_\mu J_\xi^\mu = \frac{1}{8\pi^2} \text{Tr}(G_{\mu\nu} \tilde{G}^{\mu\nu}), \quad J_\xi^\mu = f_\xi^2 \partial^\mu \xi.$$

- $U(1)_{\text{chiral}} \times U(1)_\xi$  as spurion symmetries with  $(\theta, M)$  as spurion fields:

$$U(1)_{\text{chiral}} : \begin{aligned} \Phi &\rightarrow e^{2i\alpha q_-} \Phi, \\ \theta &\rightarrow \theta + 2\alpha q_-, \\ M &\rightarrow e^{-2i\alpha q_-} M \end{aligned} \quad U(1)_\xi : \begin{aligned} \xi &\rightarrow \xi + \varepsilon, \\ \theta &\rightarrow \theta + \varepsilon. \end{aligned}$$

2 separate mass-generating terms:

$$V = V_1(\xi - i \ln \det(\Phi) - \theta) + V_2(M\Phi + M^\dagger \Phi^\dagger)$$

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with  $f \ll f_\xi$  and  $m_\xi \ll m_\eta < m_\sigma$

Viable inflationary model requires control over perturbative QM corrections:

Weinberg ('79)

- Non-renormalizable corrections have to be compatible with  $U(1)$  symmetries:

★ derivative terms:  $M_{UV}^{-4} |\partial\Phi^\dagger \partial\Phi|^2$ ,  $M_{UV}^{-2} |\Phi|^2 |\partial\Phi|^2$

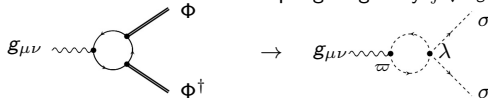
★ potential & mixed terms

$\Rightarrow$  additionally suppressed  
by powers of  $\frac{f}{f_\xi} \sim 10^{-3}$

- Loop-corrections for perturbative  $\Phi$ -interactions

★ 1-loop effective action  $V^{1-loop} \sim (-\mu^2 + 3\lambda|\Phi|^2)^2 \left\{ \ln \left( \frac{-\mu^2 + 3\lambda|\Phi|^2}{\Lambda_f^2} \right) - \frac{3}{2} \right\}$  Coleman-Weinberg ('73)  
 $\rightsquigarrow$  proper resummation using Callan-Symanzik equation for  $V_{\text{eff}}$  maintains vacuum structure

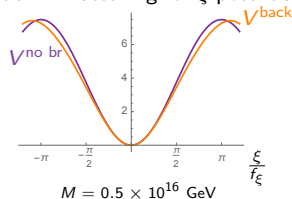
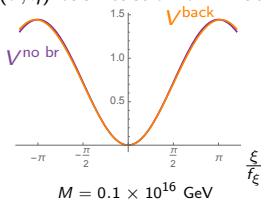
★ Induced non-minimal coupling to gravity  $\int \sqrt{-g} \varpi |\Phi|^2 R$  Hill-Salopek ('92)



Solving RGE for  $\varpi \rightarrow$  IR-fixed point  $\varpi = 0$  Voloshin-Dolgov ('82)

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- $(\sigma, \eta)$ -backreaction on inflationary potential  $\rightsquigarrow$  flattening for  $\xi$ -potential

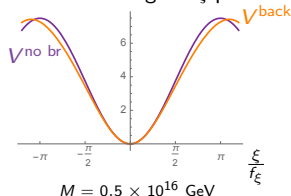
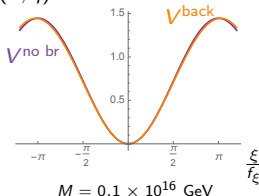


- Infladron-backreaction: hierarchy  $m_\xi \ll m_\eta < m_\sigma$  has to prevail along inflationary trajectory  $\rightsquigarrow$  OK when  $\frac{f}{f_\xi} \sim 10^{-3}$   
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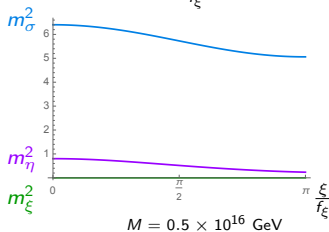
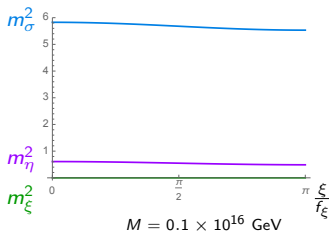


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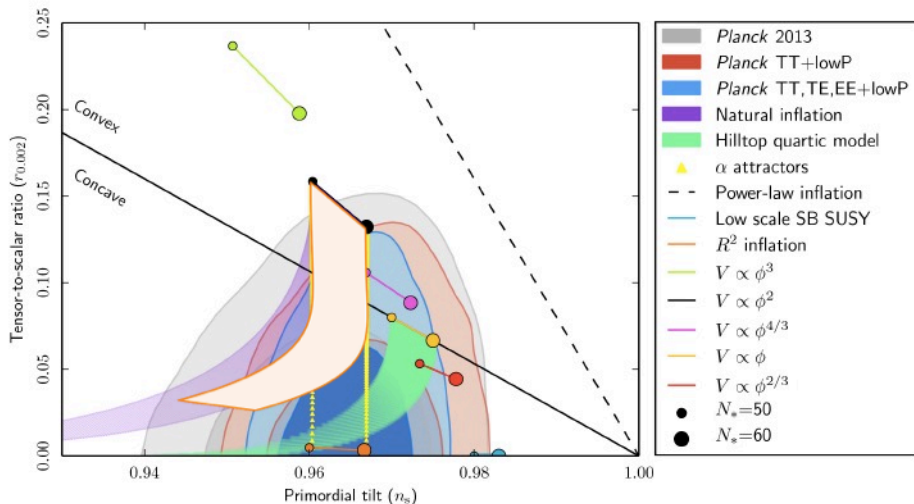


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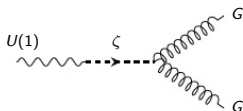


based on PLANCK A&A 594, A20 (2016)

# Consistency & 4-Fermion Couplings

w/ Shiu-Ye 1503.01015, 1503.02965 [hep-th]

- $U(1)$  gauge invariance requires presence of chiral fermions  $\psi$



- Integrating out massive  $U(1)$  boson  
 $\rightsquigarrow$  1 axion  $\xi$  + 1 non-Abelian gauge group + chiral fermions

$$\mathcal{S} = \int \frac{1}{2} d\xi \wedge \star_4 d\xi - \frac{1}{8\pi^2} \frac{\xi}{f_\xi} \text{Tr}(G \wedge G) - \frac{\mathcal{C}}{f_2^2} \underbrace{\mathcal{J}_\psi \wedge \star_4 \mathcal{J}_\psi}_{4\text{-fermion}} + \mathcal{L}_\psi$$

- Work out infrared vacuum for 1 generation  $\psi$

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The diagram shows two terms separated by a plus sign, followed by an equals zero. The first term is a tree-level exchange of a  $U(1)$  gauge boson (represented by a wavy line) between two vertices. The incoming line is labeled  $U(1)$  and the outgoing line is labeled  $\zeta$ . The two vertices are connected by two wavy lines, each labeled  $G$ . The second term is a fermion loop diagram. It consists of a triangle with a wavy line labeled  $U(1)$  on the left side. The top and bottom sides of the triangle are fermion lines labeled  $\psi$  with arrows pointing clockwise. The right side of the triangle is connected to two wavy lines, each labeled  $G$ . The entire expression is set equal to zero.

“reversed” GS mechanism

Aldazabel-Franco-Ibáñez-Rábadan-Uranga ('01)

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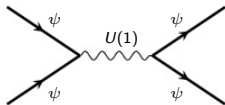
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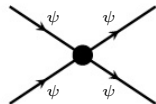
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Nambu-Jona-Lasinio ('61)

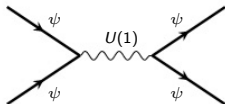
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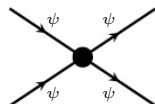
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